

Supersymmetric solutions and attractor equations in 4-dimensional supergravities

Tomás Ortín

(I.F.T. UAM/CSIC, Madrid)

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- 1 The search for **all** 4-d susy solutions
- 5 Review of the $N=2$ case
- 7 The $N = 2$ Killing Spinor Equations (KSEs)
- 9 The $N = 2$ spinor-bilinears algebra
- 10 The $N = 2$ Killing Spinor Identities (KSI)s
- 12 The $N = 2$ supersymmetric solutions
- 14 The all- N formulation of 4-d sugras
- 18 The all- N Killing Spinor Equations (KSEs)
- 19 The all- N spinor-bilinears algebra
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- 24 The all- N supersymmetric solutions
- 35 Final comments

1 – The search for all 4-d susy solutions

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Spinor-bilinears method

The **spinor** -bilinears method is specially suited for the $N = 2, d = 4$ case^a: a **Killing spinor** consists of 2 **Weyl spinors** that can be used to construct a tetrad and the complete geometry (equivalent to the **Newman -Penrose** formalism).

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For $N > 2$ there are **too many spinor bilinears** and we do not know how to extract the (**not spacetime-geometric**) information they must surely contain without breaking the symmetries of the theory.

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In this talk we are going to show how to solve those problems and determine the general form of **all** the timelike **supersymmetric** solutions of all $d = 4$ supergravities using the **spinor-bilinear method**.

One of our main results is that the timelike **supersymmetric** solutions of $N > 2, d = 4$ theories are related to those of the $N = 2, d = 4$ theories found in Hübscher, Meessen & O. (2006).

We start by reviewing them.

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The n complex scalars are encoded into the $2\bar{n}$ -dimensional symplectic section ($\bar{n} = 1 + n$)

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This is an extremely redundant (but useful) description of the scalars.

The **supersymmetry** transformations of the **fermions** are

$$\delta_{\epsilon} \psi_{I\mu} = \mathfrak{D}_{\mu} \epsilon_I + \varepsilon_{IJ} T^{+}_{\mu\nu} \gamma^{\nu} \epsilon^J,$$

$$\delta_{\epsilon} \lambda^{iI} = i \not{\partial} Z^i \epsilon^I + \varepsilon^{IJ} G^{i+} \epsilon_J.$$

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where the graviphoton and matter vector field strengths are

$$T^{+} = \langle \mathcal{V} | \mathcal{F}^{+} \rangle, \quad G^{i+} = \frac{i}{2} \mathcal{G}^{ij*} \langle \mathcal{D}_{j*} \mathcal{V}^{*} | \mathcal{F}^{+} \rangle, \quad \mathcal{F}^{+} \equiv \begin{pmatrix} F^{\Lambda+} \\ \mathcal{N}^{*}_{\Lambda\Sigma} F^{\Sigma+} \end{pmatrix},$$

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\mathfrak{D} is the **Lorentz**-, **Kähler**- and $SU(2)$ - covariant derivative (**Kähler** + $SU(2)$ = $U(2)$)

$$\mathfrak{D}_{\mu} \epsilon_I = \left(\partial_{\mu} + \frac{1}{4} \omega_{\mu}{}^{ab} \gamma_{ab} + \frac{i}{2} \mathcal{Q}_{\mu} \right) \epsilon_I + A_{\mu I}{}^J \epsilon_J,$$

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and where $U^{\alpha I}_u(q)$ is the **Quadbein**. The action for the bosonic fields is

$$S = \int d^4x \sqrt{|g|} \left[R + 2\mathcal{G}_{ij*} \partial_{\mu} Z^i \partial^{\mu} Z^{*j*} + 2\mathbb{H}_{uv} \partial_{\mu} q^u \partial^{\mu} q^v \right. \\ \left. + 2\Im \mathcal{N}_{\Lambda\Sigma} F^{\Lambda\mu\nu} F^{\Sigma}_{\mu\nu} - 2\Re \mathcal{N}_{\Lambda\Sigma} F^{\Lambda\mu\nu} \star F^{\Sigma}_{\mu\nu} \right].$$

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The goal is to find **all** the bosonic field configurations $\{e^a{}_\mu, A^\Lambda{}_\mu, Z^i, q^u\}$ such that the above **KSEs** admit at least one solution ϵ^I .

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5. Impose the independent equations of motion on the **supersymmetric** configurations we just identified.

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The 4-d Fierz identities imply that $V_a \equiv V^I_{I a}$ is always non-spacelike:

$$V^2 = -V^I_J \cdot V^J_I = 2M^{IJ} M_{IJ} = 4|X|^2 \geq 0.$$

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With them one can construct a tetrad

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with $\sigma^0 = 1$ and σ^m the 2×2 Pauli matrices as an orthonormal tetrad in which $V^0 = \sqrt{2}V$ is timelike and the V^m s are spacelike.

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with $\sigma^0 = 1$ and σ^m the 2×2 Pauli matrices as an orthonormal tetrad in which $V^0 = \sqrt{2}V$ is timelike and the V^m s are spacelike. (This will not work for $N > 2$!)

5 – The $N = 2$ Killing Spinor Identities (KSI)s

If we assume that a given **bosonic** field configuration admits a **Killing spinor** ϵ_I , then we find that the (*off-shell*) “equations of motion” $\{\mathcal{E}^{\mu\nu}, \mathcal{E}^\mu, \mathcal{E}^i, \mathcal{E}_u\}$ satisfy the **KSI**s:

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6. $\mathcal{E}_{i^*} = 2\left(\frac{X}{X^*}\right)^{1/2} \langle \mathcal{E}^0 | \mathcal{D}_{i^*} \mathcal{V}^* \rangle, (\Rightarrow \text{attractor mechanism})$

The only independent equations of motion that have to be imposed on $N = 2$, $d = 4$ supersymmetric configurations are

$$\mathcal{E}^0 = 0.$$

6 – The $N = 2$ supersymmetric solutions

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1. Define the $U(1)$ -neutral real **symplectic** vectors \mathcal{R} and \mathcal{I}

$$\mathcal{R} + i\mathcal{I} \equiv \mathcal{V}/X.$$

(\Rightarrow No **Kähler** nor $SU(2)$ **gauge** -fixing are necessary!)

6 – The $N = 2$ supersymmetric solutions

They can be constructed as follows:

1. Define the $U(1)$ -neutral real **symplectic** vectors \mathcal{R} and \mathcal{I}

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4. The **scalars** Z^i are given by the quotients

$$Z^i = \frac{\mathcal{V}^i/X}{\mathcal{V}^0/X} = \frac{\mathcal{R}^i + i\mathcal{I}^i}{\mathcal{R}^0 + i\mathcal{I}^0}.$$

5. The **hyperscalars** $q^u(x)$ are the mappings satisfying

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7. The vector field strengths are

$$\mathcal{F} = -\frac{1}{2}d(\mathcal{R}\hat{V}) - \frac{1}{2}\star(\hat{V} \wedge d\mathcal{I}), \quad \hat{V} = 2\sqrt{2}|X|^2(dt + \omega).$$

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All 4-d *supergravity* multiplets can be written in the form

$$\{e^a{}_{\mu}, \psi_{I\mu}, A^{IJ}{}_{\mu}, \chi_{IJK}, P_{IJKL\mu}, \chi^{IJKLM}\}, \quad I, J, \dots = 1, \dots, N,$$

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The price to pay for using this representation is that all the fields that can be related by $SU(N)$ **duality** relations, are:

- $N = 4$: $P^{*iIJ} = \frac{1}{2}\varepsilon^{IJKL}P_{iKL}$, and $\lambda_{iI} = \frac{1}{3!}\varepsilon_{IJKL}\lambda_i^{IJK}$.
- $N = 6$: $P^{*IJ} = \frac{1}{4!}\varepsilon^{IJK_1\dots K_4}P_{K_1\dots K_4}$, $\chi_{IJK} = \frac{1}{3!}\varepsilon_{IJKLMN}\lambda^{IJK}$,
and $\chi^{I_1\dots I_5} = \varepsilon^{I_1\dots I_5J}\lambda_J$.
- $N = 8$: $P^{*I_1\dots I_4} = \frac{1}{4!}\varepsilon^{I_1\dots I_4J_1\dots J_4}P_{J_1\dots J_4}$, and $\chi_{I_1I_2I_3} = \frac{1}{5!}\varepsilon_{I_1I_2I_3J_1\dots J_5}\chi^{J_1\dots J_5}$.

These constraints must be taken into account in the action.

The scalars are encoded into the $2\bar{n}$ -dimensional ($\bar{n} \equiv n + \frac{N(N-1)}{2}$) symplectic vectors

$$\mathcal{V}_{IJ} = \begin{pmatrix} f^{\Lambda}_{IJ} \\ h_{\Lambda IJ} \end{pmatrix}, \quad \text{and} \quad \mathcal{V}_i = \begin{pmatrix} f^{\Lambda}_i \\ h_{\Lambda i} \end{pmatrix}, \quad \Lambda = 1, \dots, \bar{n},$$

normalized

$$\langle \mathcal{V}_{IJ} | \mathcal{V}^{*KL} \rangle = -2i\delta^{KL}_{IJ}, \quad \langle \mathcal{V}_i | \mathcal{V}^{*j} \rangle = -i\delta_i^j.$$

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They can be combined into the $Usp(\bar{n}, \bar{n})$ matrix

$$U \equiv \frac{1}{\sqrt{2}} \begin{pmatrix} f + ih & f^* + ih^* \\ f - ih & f^* - ih^* \end{pmatrix}.$$

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The **graviphotons** A^{IJ}_{μ} do not appear directly, only through the “dressed” vectors

$$A^{\Lambda}_{\mu} \equiv \frac{1}{2} f^{\Lambda}_{IJ} A^{IJ}_{\mu} + f^{\Lambda}_i A^i_{\mu}.$$

The **supersymmetry** transformations of the **fermionic** fields are

$$\delta_{\epsilon}\psi_{I\mu} = \mathcal{D}_{\mu}\epsilon_I + T_{IJ}{}^{+}{}_{\mu\nu}\gamma^{\nu}\epsilon^J,$$

$$\delta_{\epsilon}\chi_{IJK} = -\frac{3i}{2} T_{[IJ}{}^{+}\epsilon_{K]} + i P_{IJKL}\epsilon^L,$$

$$\delta_{\epsilon}\lambda_{iI} = -\frac{i}{2} T_i{}^{+}\epsilon_I + i P_{iIJ}\epsilon^J,$$

$$\delta_{\epsilon}\chi_{IJKLM} = -5i P_{[IJKL}\epsilon_{M]} + \frac{i}{2}\epsilon_{IJKLMN} T^{-}\epsilon^N + \frac{i}{4}\epsilon_{IJKLMNOP} T^{NO-}\epsilon^P,$$

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$$T_{IJ}^{+} = \langle \mathcal{V}_{IJ} | \mathcal{F}^{+} \rangle, \quad T_i^{+} = \langle \mathcal{V}_i | \mathcal{F}^{+} \rangle, \quad \mathcal{F}_{\Lambda}^{+} = \mathcal{N}_{\Lambda\Sigma}^{*} F^{\Sigma+},$$

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and where

$$\mathfrak{D}_{\mu}\epsilon_I \equiv \nabla_{\mu}\epsilon_I - \epsilon_J \Omega_{\mu}{}^J{}_I,$$

and $\Omega_{\mu}{}^J{}_I$ is the pullback of the connection of the **scalar** manifold ($\subset U(N)$).

The action for the **bosonic** fields is

$$S = \int d^4x \sqrt{|g|} \left[R + 2\Im \mathcal{N}_{\Lambda\Sigma} F^{\Lambda\mu\nu} F^{\Sigma}_{\mu\nu} - 2\Re \mathcal{N}_{\Lambda\Sigma} F^{\Lambda\mu\nu} \star F^{\Sigma}_{\mu\nu} \right. \\ \left. + \frac{2}{4!} \alpha_1 P^{*IJKL}{}_{\mu} P_{IJKL}{}^{\mu} + \alpha_2 P^{*iIJ}{}_{\mu} P_{iIJ}{}^{\mu} \right],$$

where

$$\mathcal{N} = h f^{-1} = \mathcal{N}^T, \quad h_{\Lambda} = \mathcal{N}_{\Lambda\Sigma} f^{\Sigma}, \quad \mathcal{D}h_{\Lambda} = \mathcal{N}_{\Lambda\Sigma}^* \mathcal{D}f^{\Lambda}.$$

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For $N = 5$: $\mathcal{E}^{IJKL} = \mathfrak{D}^{\mu} P^{*IJKL}{}_{\mu} + 6T^{[IJ]-}{}_{\mu\nu} T^{[KL]-}{}^{\mu\nu}.$ etc.

8 – The all- N Killing Spinor Equations (KSEs)

For all values of N the independent KSEs take the form

$$\begin{aligned} \mathcal{D}_\mu \epsilon_I + T_{IJ}^+{}_{\mu\nu} \gamma^\nu \epsilon^J &= 0, \\ \mathcal{P}_{IJKL} \epsilon^L - \frac{3}{2} \mathcal{T}_{[IJ}^+ \epsilon_{K]} &= 0, \\ \mathcal{P}_{iIJ} \epsilon^J - \frac{1}{2} \mathcal{T}_i^+ \epsilon_I &= 0, \\ \mathcal{P}_{[IJKL} \epsilon_{M]} &= 0, \\ \mathcal{P}_{i[IJ} \epsilon_{K]} &= 0. \end{aligned}$$

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Again, our goal is to find **all** the bosonic field configurations $\{e^a{}_\mu, A^\Lambda{}_\mu, P_{IJKL\mu}, P_{iIJ\mu}\}$ such that the above KSEs admit at least one solution ϵ^I .

9 – The all- N spinor-bilinears algebra

The independent bilinears that we can construct with one $U(N)$ vector of Weyl spinors ϵ_I are:

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The independent bilinears that we can construct with one $U(N)$ vector of Weyl spinors ϵ_I are:

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We only consider the timelike case.
3. We can choose a tetrad $\{e^a_\mu\}$ such that $e^0_\mu \equiv \frac{1}{\sqrt{2}} |M|^{-1} V_\mu$. Then, defining $V^m_\mu \equiv |M| e^m_\mu$ we can decompose

$$V^I_{J \mu} = \frac{1}{2} \mathcal{J}^I_J V_\mu + \frac{1}{\sqrt{2}} (\sigma^m)^I_J V^m_\mu,$$

where $\mathcal{J}^I_J = 2M^{IK} M_{JK} |M|^{-2}$ is a rank 2 projector (Tod):

$$\mathcal{J}^2 = \mathcal{J}, \quad \mathcal{J}^I_I = +2, \quad \mathcal{J}^I_J \epsilon^J = \epsilon^I.$$

The main properties satisfied by the three σ^m matrices are:

$$\sigma^m \sigma^n = \delta^{mn} \mathcal{J} + i\varepsilon^{mnp} \sigma^p,$$

$$\mathcal{J} \sigma^m = \sigma^m \mathcal{J} = \sigma^m,$$

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$$\mathcal{J}^K{}_J \mathcal{J}^L{}_I = \frac{1}{2} \mathcal{J}^K{}_I \mathcal{J}^L{}_J + \frac{1}{2} (\sigma^m)^K{}_I (\sigma^m)^L{}_J,$$

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$\{\mathcal{J}, \sigma^1, \sigma^2, \sigma^3\}$ is an x -dependent basis of a $\mathfrak{u}(2)$ subalgebra of $\mathfrak{u}(N)$ in the 2-dimensional eigenspace of \mathcal{J} of eigenvalue +1 and provide a basis in the space of Hermitean matrices A satisfying $\mathcal{J} A \mathcal{J} = A$

10 – The all-N Killing Spinor Identities (KSIs)

If we assume that a given **bosonic** field configuration admits a **Killing spinor** ϵ_I , then we find that the (*off-shell*) “equations of motion” $\{\mathcal{E}^{\mu\nu}, \mathcal{E}^\mu, \mathcal{E}^{IJKL}, \mathcal{E}^{iIJ}\}$ satisfy the **KSIs** ($\tilde{\mathcal{J}}^I{}_J \equiv \delta^I{}_J - \mathcal{J}^I{}_J$):

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4. $\mathcal{E}^{00} = -2\sqrt{2} \langle \mathcal{E}^0 \mid \Re \left(\nu_{IJ} \frac{M^{IJ}}{|M|} \right) \rangle, \text{ (Bogomol'nyi bound)}$

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$N = 4$: $\left\{ \begin{array}{l} \mathcal{E}^{IJKL} = -2\sqrt{2} \frac{M^{[IJ}}{|M|} \langle \mathcal{E}^0 \mid \mathcal{V}^{*|KL]} \rangle, \\ \mathcal{E}_{iIJ} = -2\sqrt{2} \left\{ \frac{M_{IJ}}{|M|} \langle \mathcal{E}^0 \mid \mathcal{V}_i \rangle + \frac{1}{2} \epsilon_{IJKL} \frac{M^{KL}}{|M|} \langle \mathcal{E}^0 \mid \mathcal{V}^{*i} \rangle \right\}, \end{array} \right.$

etc.

The only independent equations of motion that have to be imposed on **any** $d = 4$ supersymmetric configuration are

$$\mathcal{E}^0 = 0 .$$

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We also have to impose the constraint

$$\mathcal{J} d\sigma^m \mathcal{J} = 0.$$

Once the $U(2)$ subgroup has been chosen, we can split the Vielbeins $P_{IJKL\mu}$ and $P_{iIJ\mu}$, into associated to the would-be **vector multiplets** in the $N = 2$ **truncation**

$$P_{IJKL} \mathcal{J}^I_{[M} \mathcal{J}^J_N \tilde{\mathcal{J}}^K_P \tilde{\mathcal{J}}^L_{Q]}, \quad \text{and} \quad P_{iIJ} \mathcal{J}^I_{[K} \mathcal{J}^J_{L]},$$

which are driven by the *attractor mechanism* (*i.e.* they are determined by the **electric** and **magnetic** charges) and those associated to the **hypermultiplets**

$$P_{IJKL} \mathcal{J}^I_{[M} \tilde{\mathcal{J}}^J_N \tilde{\mathcal{J}}^K_P \tilde{\mathcal{J}}^L_{Q]}, \quad \text{and} \quad P_{iIJ} \mathcal{J}^I_{[K} \tilde{\mathcal{J}}^J_{L]}.$$

which are not.

In **hyper**-less solutions (*e.g.* black holes) the σ^m s matrices are not needed at all.

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where

$$|M|^{-2} = (M^{IJ} M_{IJ})^{-2} = \langle \mathcal{R} | \mathcal{I} \rangle ,$$

$$(d\omega)_{mn} = 2\epsilon_{mnp} \langle \mathcal{I} | \partial^p \mathcal{I} \rangle .$$

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$$F = -\frac{1}{2}d(\mathcal{R}\hat{V}) - \frac{1}{2}\star(\hat{V} \wedge d\mathcal{I}), \quad \hat{V} = \sqrt{2}|M|^2(dt + \omega).$$

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can be found from \mathcal{R} and \mathcal{I} , while those in the **hypers** must be found independently by solving

$$P_{IJKLm} \mathcal{J}^I_{[M} \tilde{\mathcal{J}}^J_N \tilde{\mathcal{J}}^K_P \tilde{\mathcal{J}}^L_{Q]} (\sigma^m)^Q_R = 0,$$

$$P_{iIJm} \mathcal{J}^I_{[K} \tilde{\mathcal{J}}^J_{L]} (\sigma^m)^L_M = 0,$$

which solve their equations of motion according to the *Killing Spinor Identities*.

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A simple derivation of the attractor flow eqs. in $N = 1, d = 5$ supergravity

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for some coordinate ρ . Let's define the **central charge**

$$\mathcal{Z}[\phi(\rho), q] \equiv h^I(\phi) q_I.$$

Then, using $h^I h_I = 1$ and $dh^I h_I = h^I dh_I = 0$

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We define the central charges

$$\mathcal{Z}_{IJ}[\phi(\rho), q] \equiv \langle \mathcal{V}_{IJ} | q \rangle = p^\Lambda h_{\Lambda IJ} - q_\Lambda f^\Lambda_{IJ},$$

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Then

$$\begin{aligned} \mathfrak{D} \frac{M^{IJ}}{|M|^2} &= \mathfrak{D} \left(\frac{M^{KL}}{|M|^2} \frac{i}{2} \langle \mathcal{V}_{KL} | \mathcal{V}^{*IJ} \rangle \right) = \frac{i}{2} \mathfrak{D} \langle (\mathcal{R} + i\mathcal{I}) | \mathcal{V}^{*IJ} \rangle \\ &= \frac{i}{2} \langle d(\mathcal{R} + i\mathcal{I}) | \mathcal{V}^{*IJ} \rangle = \frac{i}{2} \langle d(\mathcal{R} - i\mathcal{I}) | \mathcal{V}^{*IJ} \rangle - \langle d\mathcal{I} | \mathcal{V}^{*IJ} \rangle \\ &= \frac{i}{2} \frac{M_{KL}}{|M|^2} \langle d\mathcal{V}^{*KL} | \mathcal{V}^{*IJ} \rangle - \langle q | \mathcal{V}^{*IJ} \rangle d\rho \\ &= \frac{1}{2} P^*{}_{KLIJ} \frac{M_{KL}}{|M|^2} + \mathcal{Z}^{*IJ}[\phi(\rho), q] d\rho. \end{aligned}$$

With the above identity we can compute

$$d|M|^{-2} = \frac{M_{IJ}}{|M|^2} \mathfrak{D} \frac{M^{IJ}}{|M|^2} + \frac{M^{IJ}}{|M|^2} \mathfrak{D} \frac{M_{IJ}}{|M|^2} = \frac{M_{IJ} \mathcal{Z}^{*IJ} + M^{IJ} \mathcal{Z}_{IJ}}{|M|^2} d\rho,$$

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which leads to the flow equation ($N \geq 4$)

$$P^{*MN[IJ} \mathcal{J}^K{}_M \mathcal{J}^L]_N = -M^{[IJ} \mathcal{Z}^{*KL]} d\rho.$$

The third flow equation ($N = 2, 3, 4, 6$) follows from

$$\begin{aligned} \frac{1}{2} \frac{M^{IJ}}{|M|^2} P_{iIJ} &= -\frac{i}{2} \frac{M^{IJ}}{|M|^2} \langle d\mathcal{V}_{IJ} \mid \mathcal{V}_i \rangle = -\frac{i}{2} \langle d(\mathcal{R} + i\mathcal{I}) \mid \mathcal{V}_i \rangle \\ &= \langle d\mathcal{I} \mid \mathcal{V}_i \rangle - \frac{i}{2} \langle d(\mathcal{R} - i\mathcal{I}) \mid \mathcal{V}_i \rangle \\ &= -\mathcal{Z}_i d\rho, \end{aligned}$$

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 &= -\mathcal{Z}_i d\rho,
 \end{aligned}$$

and takes the final form

$$P_{iKL} \mathcal{J}^K_I \mathcal{J}^L_J = -2M_{IJ} \mathcal{Z}_i d\rho.$$

Summarizing, we have obtained 3 differential equations:

$$\left\{ \begin{array}{l} \frac{d}{d\rho} |M|^{-1} = \Re \left(\frac{M^{IJ} \mathcal{Z}_{IJ}}{|M|} \right), \\ P^* MN [IJ \mathcal{J}^K_M \mathcal{J}^L]_N = -M^{[IJ} \mathcal{Z}^*{}^{KL]} d\rho, \\ P_{iKL} \mathcal{J}^K_I \mathcal{J}^L_J = -2M_{IJ} \mathcal{Z}_i d\rho, \end{array} \right.$$

which govern the metric function $g_{tt} = |M|^2$ and the scalars that would fit into $N = 2$ vector **supermultiplets**. The **supersymmetric attractors** are at the solutions of

$$M^{[IJ} \mathcal{Z}^*{}^{KL]} = 0, \quad (\text{for scalars in the supergravity multiplet})$$

$$\mathcal{Z}_i = 0, \quad (\text{for scalars in vector multiplets as in } N = 2)$$

Only in the $N = 2$ case the last equation is a differential equation.

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- ★ “1-line” derivations of the **attractor flow equations** can be readily found.
- ★ Much work remains to be done in order to make explicit the construction of the solutions. In particular one has to find general parametrizations of the matrices M^{IJ} and $\mathcal{J}^I{}_J$, solve the *stabilization equations*, impose the covariant constancy of \mathcal{J} etc. (Meessen & O., work in progress).