

# The tensor hierarchy and supersymmetric domain walls of $N=1, d=4$ supergravity

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Work done in collaboration with *E. Bergshoeff, O. Hohm (U. Groningen) J. Hartong (U. Bern) M. Hübscher and P. Meessen (IFT UAM/CSIC, Madrid)*

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3. The *embedding tensor method* (**Cordaro, Fré, Gualtieri, Termonia & Trigiante, arXiv:hep-th/9804056.** ) can be used to construct systematically the most general **gauged supergravities** . This construction requires the introduction of additional  $(p + 1)$ -form potentials.



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We are going to use the **embedding tensor** method to find all the  $(p + 1)$ -form potentials and the corresponding **democratic formulations** of any 4-dimensional FT with **gauge** symmetry and we are going to apply the general results to the particular case of  **$N = 1$  supergravity** .

What we are going to do in this seminar:

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5. We will use some of the new  $(p + 1)$ -form potentials to construct **supersymmetric  $p$ -brane** effective actions and solutions with sources of  **$N = 1$  supergravity**.

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## 2 – The embedding tensor method: electric gaugings

Consider a general ( $N = 1$  supergravity -inspired) 4-dimensional ungauged FT with bosonic fields  $\{Z^i, A^\Lambda\}$  (the metric plays no relevant role here)

$$S_u[Z^i, A^\Lambda] = \int \left\{ -2\mathcal{G}_{ij^*} dZ^i \wedge \star dZ^{*j^*} - 2\Im f_{\Lambda\Sigma} F^\Lambda \wedge \star F^\Sigma + 2\Re f_{\Lambda\Sigma} F^\Lambda \wedge F^\Sigma - \star V_u(Z, Z^*) \right\}.$$

with  $F^\Lambda \equiv dA^\Lambda$ , the fundamental (electric) field strengths and  $f_{\Lambda\Sigma}(Z)$ .



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The action is invariant under the local Abelian transformations

$$\delta_\Lambda A^\Sigma = d\Lambda^\Sigma.$$

Let us assume this action is invariant under the global transformations

$$\delta_\alpha Z^i = \alpha^A k_A^i(Z),$$

$$\delta_\alpha f_{\Lambda\Sigma} \equiv -\alpha^A \mathcal{L}_A f_{\Lambda\Sigma} = \alpha^A [T_{A\Lambda\Sigma} - 2T_{A(\Lambda} \Omega^\Omega f_{\Sigma)\Omega}],$$

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Gauging the global symmetries of a FT with constant parameters  $\alpha^A$  means modifying it to make it invariant when the  $\alpha^A$  are arbitrary functions  $\alpha^A(x)$ .

## *The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity*

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Each embedding tensor  $\vartheta_\Lambda^A$  defines a possible identification:

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covariant under

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## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

This only works if  $\vartheta_\Lambda^A$  is an invariant tensor

$$\delta_\Lambda \vartheta_\Sigma^A = -\Lambda^\Omega Q_{\Omega\Sigma}^A = 0, \quad Q_{\Sigma\Lambda}^A \equiv \vartheta_\Sigma^B T_{B\Lambda}{}^\Omega \vartheta_\Omega^A - \vartheta_\Sigma^B \vartheta_\Lambda^C f_{BC}^A.$$

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It is customary to define the generators

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which satisfy the algebra

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Then we construct the covariant 2-form field strengths

$$F^\Lambda = dA^\Lambda + \frac{1}{2} X_{\Sigma\Omega}{}^\Lambda A^\Sigma \wedge A^\Omega,$$

and the *gauge* -invariant action of the *electrically gauged* FT takes the form

$$S_{\text{eg}}[Z^i, A^\Lambda] = \int \left\{ -2\mathcal{G}_{ij}{}^* \mathcal{D}Z^i \wedge \star \mathcal{D}Z^{*j} - 2\Im f_{\Lambda\Sigma} F^\Lambda \wedge \star F^\Sigma + 2\Re f_{\Lambda\Sigma} F^\Lambda \wedge F^\Sigma - \star V_{\text{eg}}(Z, Z^*) \right\}$$

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$\Rightarrow$  One can define magnetic (dual) 1-forms  $A_\Lambda$  which one may use as gauge fields: if the Maxwell equations are

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$\Rightarrow$  The theory (equations of motion) has more **non-perturbative global** symmetries that can be **gauged**. They include **electric -magnetic duality** rotations:

$$\delta_\alpha Z^i = \alpha^A k_A^i(Z),$$

$$\delta_\alpha f_{\Lambda\Sigma} = \alpha^A \{-T_{A\Lambda\Sigma} + 2T_{A(\Lambda}{}^\Omega f_{\Sigma)\Omega} - T_A{}^{\Omega\Gamma} f_{\Omega\Lambda} f_{\Gamma\Sigma}\},$$

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Now we need to relate the  $\alpha^A$  to the gauge parameters of the 1-forms  $\Lambda^\Lambda$  or  $\Lambda_\Lambda$ . We need new (magnetic) components for the embedding tensor:  $\vartheta^{\Lambda A}$ . Then

$$\alpha^A(x) \equiv \Lambda^\Sigma \vartheta_{\Sigma}^A + \Lambda_\Sigma \vartheta^{\Sigma A}, \quad A^A \equiv A^\Sigma \vartheta_{\Sigma}^A + A_\Sigma \vartheta^{\Sigma A}.$$

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Knowing (**Gaillard & Zumino**) that the  $T_A$  matrices either belong to  $\mathfrak{sp}(2n_V, \mathbb{R})$  or vanish, we introduce the **symplectic** notation

$$A^M \equiv \begin{pmatrix} A^\Sigma \\ A_\Sigma \end{pmatrix} \quad \vartheta_M^A \equiv (\vartheta_{\Sigma}^A, \vartheta^{\Sigma A}) \quad \alpha^A(x) \equiv \Lambda^M \vartheta_M^A,$$

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The **electric** and **magnetic** charges must be mutually **local** (**de Wit, Samtleben & Trigiante**, arXiv:hep-th/0507289):

$$Q^{AB} \equiv \frac{1}{4} \vartheta^{MA} \vartheta_M^B = 0.$$



Now we can repeat the procedure of the **electric** case:

First we construct derivatives  $\mathfrak{D}$

$$\mathfrak{D}Z^i \equiv dZ^i + A^M \vartheta_M^A k_A^i,$$

covariant under

$$\delta_\Lambda Z^i = \Lambda^M \vartheta_M^A k_A^i(Z),$$

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which only works if  $\vartheta_M^A$  is an invariant tensor

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Before moving forward, we must impose another constraint on the **embedding tensor** on top of the two quadratic ones  $Q_{MN}^A = Q^{AB} = 0$ :

$$L_{MNP} \equiv X_{(MNP)} = \vartheta_{(M}^A T_{ANP)} = 0.$$

This *linear* or *representation constraint* is based on **supergravity** and eliminates certain possible representations of the **embedding tensor**. On the other hand, we cannot construct **gauge**-covariant 2-form field strengths  $F^M$  without it!

## 4 – The 4-d tensor hierarchy

To construct the **gauge** -covariant 2-form field strengths  $F^M$  we take the covariant derivative of the **gauge** -covariant “field strength”  $\mathcal{D}Z^i$ :

$$\mathcal{D}\mathcal{D}Z^i = [dA^M + \frac{1}{2}X_{NP}{}^M A^N \wedge A^P] \vartheta_M{}^A k_A{}^i,$$

which suggests the definition

$$F^M \equiv dA^M + \frac{1}{2}X_{NP}{}^M A^N \wedge A^P + \Delta F^M, \quad \vartheta_M{}^A \Delta F^M = 0,$$

so we have the **Bianchi** identity

$$\mathcal{D}\mathcal{D}Z^i = F^M \vartheta_M{}^A k_A{}^i .$$

## 4 – The 4-d tensor hierarchy

To construct the **gauge** -covariant 2-form field strengths  $F^M$  we take the covariant derivative of the **gauge** -covariant “field strength”  $\mathcal{D}Z^i$ :

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To determine  $\Delta G_C{}^M$  we need to find invariant tensors that vanish upon contraction with  $Y_{AM}{}^C$ . They appear automatically when we take the **gauge** -covariant derivative of the **Bianchi** identity and  $G_C{}^M$  (if we “forget” we are in 4 dimensions!).

Acting with  $\mathfrak{D}$  on the **Bianchi** identity of  $H_A$  we find

$$Y_{AM}{}^C \{ \mathfrak{D}G_C{}^M - F^M \wedge H_A \} = 0, \Rightarrow \mathfrak{D}G_C{}^M = F^M \wedge H_A + \Delta \mathfrak{D}G_C{}^M,$$

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This implies that there are 3 such tensors  $W_C{}^{MAB}$ ,  $W_{CNPQ}{}^M$ ,  $W_{CNP}{}^{EM}$  that vanish contracted with  $Y_{AM}{}^C$  and which we can use to build  $\Delta G_C{}^M$ .



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The natural solution is

$$\Delta G_C{}^M = W_C{}^{MAB} D_{AB} + W_{CNPQ}{}^M D^{NPQ} + W_{CNP}{}^{EM} D_E{}^{NP},$$

and  $\delta_\Lambda D_{AB}$ ,  $\delta_\Lambda D^{NPQ}$ ,  $\delta_\Lambda D_E{}^{NP}$  will follow from the gauge-covariance of  $G_C{}^M$ .

What have we got so far just by asking for covariance under **gauge** transformations?

## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

What have we got so far just by asking for covariance under gauge transformations?

⇒ A tower of  $(p + 1)$ -forms  $A^M, B_A, C_C^M, D_{AB}, D^{NPQ}, D_E^{NP}$  related by gauge transformations.

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**But, what does it mean?**  
**What is the meaning of the additional fields?**

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- These two **duality** relations together with the **Bianchi** identity  $\mathcal{D}F^M = Z^{MA} H_A$  give a set of **electric -magnetic duality** -covariant **Maxwell** equations:

$$\mathcal{D}F^\Lambda = -\frac{1}{4} \vartheta_\Lambda^A \star j_A , \quad \mathcal{D}G_\Lambda = \frac{1}{4} \vartheta^\Lambda A \star j_A .$$

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→ The 3-forms  $C_C^M$  must be dual to constants: the embedding tensor  $\vartheta_M^C$ . This duality is expressed through the formula

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→ Using the three duality relations in the Bianchi identity of  $H_A$  we get

$$\mathcal{D} \star j_A = 4T_{AMN} G^M \wedge G^N + \star Y_A^{MC} \frac{\partial V}{\partial \vartheta_M^C} .$$

→ The 3-forms  $C_C^M$  must be **dual** to constants: the **embedding tensor**  $\vartheta_M^C$ . This **duality** is expressed through the formula

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→ Using the three **duality** relations in the **Bianchi** identity of  $H_A$  we get

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This equation is similar to the consistency condition (**gauge** or **Noether** identity) that **Noether** currents must satisfy off-shell in FTs with **gauge** invariance:

$$\mathcal{D} \star j_A = -2(k_A^i \mathcal{E}_i + \text{c.c.}) + 4T_{AMN} G^M \wedge G^N + \star Y_A^{MC} \frac{\partial V}{\partial \vartheta_M^C} ,$$

where  $\mathcal{E}_i$  is the e.o.m. of  $Z^i$ . Both equations, together, imply

$$k_A^i \mathcal{E}_i + \text{c.c.} = 0 ,$$

which is equivalent to the scalar e.o.m. for symmetric  $\sigma$ -models.



## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

✎ Finally, the indices of the 3 4-forms  $D_{AB}$ ,  $D^{NPQ}$ ,  $D_E^{NP}$  are conjugate to those of the constraints  $Q^{AB}$ ,  $Q_{NPQ}$ ,  $Q_{NP}^E$ . They are Lagrange multipliers enforcing them.

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To confirm this interpretation we must construct a **gauge** -invariant action for **all** these fields, including the **embedding tensor**  $\vartheta_M^A(x)$ .

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This **gauge** -invariant action is given by

$$\begin{aligned}
 S[g_{\mu\nu}, Z^i, A^M, B_A, C_A^M, D_E^{NP}, D_{AB}, D^{MNP}, \vartheta_M^A] = & \\
 \int \{ & -2\mathcal{G}_{ij^*} \mathcal{D}Z^i \wedge \star \mathcal{D}Z^{*j^*} + 2F^\Sigma \wedge G_\Sigma - \star V \\
 & -4Z^{\Sigma A} B_A \wedge (F_\Sigma - \frac{1}{2} Z_\Sigma^B B_B) - \frac{4}{3} X_{[MN]\Sigma} A^M \wedge A^N \wedge (F^\Sigma - Z^{\Sigma B} B_B) \\
 & -\frac{2}{3} X_{[MN]}^\Sigma A^M \wedge A^N \wedge (dA_\Sigma - \frac{1}{4} X_{[PQ]\Sigma} A^P \wedge A^Q) \\
 & -2\mathcal{D}\vartheta_M^A \wedge (C_A^M + A^M \wedge B_A) \\
 & +2Q_{NP}^E (D_E^{NP} - \frac{1}{2} A^N \wedge A^P \wedge B_E) + 2Q^{AB} D_{AB} + 2L_{MNP} D^{MNP} \} .
 \end{aligned}$$

## 6 – Application: general gaugings of $N = 1, d = 4$ supergravity

Now we want to apply our results to **gauge  $N = 1$   $d = 4$  supergravity** with generic matter content and couplings.

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- ➡ In  $N = 1$   $N = 2$  supergravity one can write  $G = G_{\text{bos}} \times H_{\text{aut}}$ , i.e.  $R$ -symmetry only acts on the fermions, which have been ignored in the construction of the tensor hierarchy.

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- ➡ In  $N = 1$   $N = 2$  **supergravity** one can write  $G = G_{\text{bos}} \times H_{\text{aut}}$ , i.e. **R-symmetry** only acts on the **fermions**, which have been ignored in the construction of the **tensor hierarchy**.

We are going to review **ungauged**  $N = 1$  **supergravity** and its **global** symmetries and then we are going to **gauge** them using the **embedding tensor** formalism.

**7 – Ungauged  $N = 1, d = 4$  supergravity**



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All **fermions** are represented by chiral 4-component spinors:

$$\gamma_5 \psi_\mu = -\psi_\mu, \text{ etc.}$$

The couplings



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The **spinors** transform as *sections* of the bundle: under **Kähler** transformations

$$\delta_\lambda \mathcal{K} = \lambda(Z) + \lambda^*(Z^*) , \quad \delta_\lambda \psi_\mu = -\frac{1}{4} [\lambda(Z) - \lambda^*(Z^*)] \psi_\mu ,$$

and their covariant derivatives contain the pullback of the **Kähler** connection 1-form  $\mathcal{Q} \equiv \mathcal{Q}_i dZ^i + \mathcal{Q}_{i^*} dZ^{*i^*}$  e.g.

$$\mathcal{D}_\mu \psi_\nu = \{ \nabla_\mu + \frac{i}{2} \mathcal{Q}_\mu \} \psi_\nu .$$

## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

$N = 1$  supergravity allows for an arbitrary holomorphic kinetic matrix  $f_{\Lambda\Sigma}(Z)$  for the vector fields which occurs in the action in the terms

$$-2\Im f_{\Lambda\Sigma} F^\Lambda \wedge \star F^\Sigma + 2\Re f_{\Lambda\Sigma} F^\Lambda \wedge F^\Sigma, \quad F^\Lambda \equiv dA^\Lambda.$$

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Finally, ungauged  $N = 1$  supergravity allows for a holomorphic superpotential  $\mathcal{W}(Z)$  which appears through the covariantly holomorphic section of Kähler weight  $(1, -1)$   $\mathcal{L}(Z, Z^*)$ :

$$\mathcal{L}(Z, Z^*) = \mathcal{W}(Z)e^{\kappa/2}, \quad \mathcal{D}_{i^*}\mathcal{L} = 0,$$

which couples to the fermions in various ways and gives rise to the scalar potential

$$V_u(Z, Z^*) = -24|\mathcal{L}|^2 + 8g^{ij^*}\mathcal{D}_i\mathcal{L}\mathcal{D}_{j^*}\mathcal{L}^*.$$

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The bosonic action is

$$\begin{aligned} S_u[g_{\mu\nu}, Z^i, A^\Lambda] &= \int \{ \star R - 2\mathcal{G}_{ij^*} dZ^i \wedge \star dZ^{*j^*} - 2\Im f_{\Lambda\Sigma} F^\Lambda \wedge \star F^\Sigma \\ &\quad + 2\Re f_{\Lambda\Sigma} F^\Lambda \wedge F^\Sigma - \star V_u(Z, Z^*) \}. \end{aligned}$$

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Main difference with the general case: the existence of  $H_{\text{aut}} = U(1)_R$ .



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- ➡ The **superpotential**  $\mathcal{L}(Z, Z^*)$  is not a fundamental field and this phase change is not a symmetry unless it can be reabsorbed into a transformation of the scalars.
- ➡ But this would mean that we are dealing with a  $A = \mathbf{a}$  symmetry and we can say that a non-vanishing superpotential breaks  $U(1)_R$  and we cannot gauge it.

## 8 – Gauging $N = 1, d = 4$ Supergravity

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➡ Then, the **spinors** ' covariant derivatives take the form

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- ➔ According to the previous discussion, the symmetries  $A = \underline{a}, \#$  are broken and cannot be **gauged** if  $\mathcal{L} \neq 0$ .

$$\mathcal{L} \neq 0 , \Rightarrow \vartheta_M{}^A (\delta_{A\underline{a}} \mathcal{P}_{\underline{a}} + \delta_{A\#} \mathcal{P}_{\#}) = 0 .$$

## 9 – The $N = 1, d = 4$ bosonic tensor hierarchy

We have found that, for non-vanishing **superpotential**, the **embedding tensor** must satisfy another constraint

$$Q_M \equiv \vartheta_M^A (\delta_A^{\underline{a}} \mathcal{P}_{\underline{a}} + \delta_A^{\#} \mathcal{P}_{\#}) = 0,$$

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→ Now ( $\mathcal{L} \neq 0$ ) the constraint  $Z^{MA} \Delta H_A = 0$  can be solved in a more general form:

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We have found that, for non-vanishing **superpotential**, the **embedding tensor** must satisfy another constraint

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This will happen in  $N = 1$  **supergravity** if we find new **Stückelberg** shifts

$$\delta' B_A \sim \delta_h B_A + Y_A \Lambda \quad \text{and} \quad \delta' C_C^M = \delta_h C_C^M + Y_C \Lambda^M.$$

## 10 – The $N = 1, d = 4$ supersymmetric tensor hierarchy

As a first step to include the **tensor hierarchy** fields into  $N = 1$  supergravity we are going to construct **supersymmetry** transformation rules such that the **local supersymmetry** algebra, to leading order in **fermions**, closes on the new fields up to **duality** relations.

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This construction requires new **duality** rules for the **supersymmetric** partners.

Observe that we are going to obtain, independently, the **gauge** transformations of the fields and will confirm or refute the hierarchy's results.

The scalars  $Z^i$

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$$\delta_{\eta} \chi^i = i \mathfrak{D} Z^i \eta^* + 2 \mathcal{G}^{ij*} \mathcal{D}_{j^*} \mathcal{L}^* \eta, \quad \mathfrak{D} Z^i = dZ^i + A^M \vartheta_M^A k_A^i.$$

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We find the expected result

$$[\delta_{\eta}, \delta_{\epsilon}] Z^i = \delta_{\text{g.c.t.}} Z^i + \delta_h Z^i,$$

$$\delta_{\text{g.c.t.}} Z^i = \mathcal{L}_{\xi} Z^i = + \xi^{\mu} \partial_{\mu} Z^i,$$

$$\delta_h Z^i = \Lambda^M \vartheta_M^A k_A^i,$$

$$\xi^{\mu} \equiv \frac{i}{4} (\bar{\epsilon} \gamma^{\mu} \eta^* - \bar{\eta} \gamma^{\mu} \epsilon^*),$$

$$\Lambda^M \equiv \xi^{\mu} A^M_{\mu}.$$

The 1-forms  $A^M$

We introduce **supersymmetric** partners  $\lambda_\Sigma$  for the **magnetic** 1-forms  $A_\Sigma$  and make the **symplectic** -covariant Ansatz

$$\begin{aligned}\delta_\epsilon A^M{}_\mu &= -\frac{i}{8} \bar{\epsilon}^* \gamma_\mu \lambda^M + \text{c.c.}, \\ \delta_\epsilon \lambda^M &= \frac{1}{2} [F^{M+} + i\mathcal{D}^M] \epsilon,\end{aligned}$$

where we have defined the **symplectic** vector

$$\mathcal{D}^M \equiv \begin{pmatrix} \mathcal{D}^\Lambda \\ \mathcal{D}_\Lambda \end{pmatrix} \equiv \begin{pmatrix} \mathcal{D}_\Lambda \\ f_{\Lambda\Sigma} \mathcal{D}^\Sigma \end{pmatrix}, \quad \mathcal{D}^\Lambda = -\Im f^{\Lambda\Sigma} (\vartheta_\Sigma^A + f_{\Sigma\Omega}^* \vartheta^{\Omega A}) \mathcal{P}_A.$$

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$$[\delta_\eta, \delta_\epsilon] A^M = \delta_{\text{g.c.t.}} A^M + \delta_h A^M,$$

where

$$\Lambda_A \equiv -T_{AMN} A^N \Lambda^M + b_A - \mathcal{P}_A \xi, \quad b_{A\mu} \equiv B_{A\mu\nu} \xi^\nu.$$



The 2-forms  $B_A$

We introduce the **supersymmetric** partners  $\zeta_A, \varphi_A$  (linear **supermultiplets** )

$$\delta_\epsilon \zeta_A = -i \left[ \frac{1}{12} H'_A + \mathcal{D} \varphi_A \right] \epsilon^* - 4 \delta_A^{\mathbf{a}} \varphi_{\mathbf{a}} \mathcal{L}^* \epsilon,$$

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which **proves** the existence of an extra **Stückelberg** shift in  $B_A$ .

The 3-forms  $C_A^M$

In this case we do not introduce any **supersymmetric** partners. We just make the Ansatz

$$\delta_\epsilon C_A^M{}_{\mu\nu\rho} = -\frac{i}{8} [\mathcal{P}_A \bar{\epsilon}^* \gamma_{\mu\nu\rho} \lambda^M - \text{c.c.}] - 3B_A{}_{[\mu\nu} \delta_\epsilon A^M{}_{|\rho]} - 2T_{APQ} A^M{}_{[\mu} A^P{}_{\nu} \delta_\epsilon A^Q{}_{|\rho]} .$$

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This corresponds to a scalar potential of the form

$$V_{e-mg} = V_u - \frac{1}{2} \Re \mathcal{D}^M \vartheta_M^A \mathcal{P}_A = V_u + \frac{1}{2} \mathcal{M}^{MN} \vartheta_M^A \vartheta_N^B \mathcal{P}_A \mathcal{P}_B,$$

where

$$\left( \mathcal{M}^{MN} \right) \equiv \begin{pmatrix} I^{\Lambda\Sigma} & I^{\Lambda\Omega} R_{\Omega\Sigma} \\ R_{\Lambda\Omega} I^{\Omega\Sigma} & I_{\Lambda\Sigma} + R_{\Lambda\Omega} I^{\Omega\Gamma} R_{\Gamma\Sigma} \end{pmatrix}, \quad \begin{aligned} f_{\Lambda\Sigma} &\equiv R_{\Lambda\Sigma} + iI_{\Lambda\Sigma}, \\ I^{\Lambda\Omega} I_{\Omega\Sigma} &\equiv \delta^\Lambda{}_\Sigma, \end{aligned}$$

so it is manifestly **symplectic** -invariant, as it must.

The 3-forms  $C, C'$

The consistency of the previous results requires the existence of a 3-form  $C$  transforming under the extra Stückelberg shift of  $B_A$ .

$$\delta_\epsilon C_{\mu\nu\rho} = -3ig\mathcal{L}\bar{\epsilon}^* \gamma_{[\mu\nu}\psi^*_{\rho]} - \frac{1}{2}\mathbf{g}\mathcal{D}_i\mathcal{L}\bar{\epsilon}^* \gamma_{\mu\nu\rho}\chi^i + \text{c.c.},$$

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If we rescale the **superpotential** by  $\mathcal{L} \rightarrow \mathbf{g}\mathcal{L}$ , the above **duality** relation takes the standard form

$$G' = \frac{1}{2}\star\frac{\partial V_{e-mg}}{\partial\mathbf{g}},$$

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So, what is the 3-form  $C$  dual to?

The 4-forms  $D_{AB}, D^{NPQ}, D_E^{NP}, D^M$

The calculations become horribly complicated and we only check the closure of the **local supersymmetry** algebra in the **ungauged**  $\vartheta_M^A = 0$  case when there are no symmetries acting on the 1-forms i.e.  $T_{AM}^N = 0$ .

The **supersymmetry** transformations are

$$\delta_\epsilon D_{AB} = -\frac{i}{2} \star \mathcal{P}_{[A} \partial_i \mathcal{P}_{B]} \bar{\epsilon} \chi^i + \text{c.c.} - B_{[A} \wedge \delta_\epsilon B_{B]},$$

$$\delta_\epsilon D^{NPQ} = 10 A^{(N} \wedge F^P \wedge \delta_\epsilon A^{Q)},$$

$$\delta_\epsilon D_E^{NP} = C_E^P \wedge \delta_\epsilon A^N.$$

$$\delta_\epsilon D^M = -\frac{i}{2} \star \mathcal{L}^* \bar{\epsilon} \lambda^M + \text{c.c.} + C \wedge \delta_\epsilon A^M.$$

This proves that  $D^M$  can be consistently added to the **supersymmetric** theory. Its role in the action will be that of **Lagrange** multiplier of the constraint  $Q_M$ .

## 11 – The supersymmetric objects of $N = 1$ supergravity



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This agrees with the results on the classification of **supersymmetric** solutions of  $N = 1$  **supergravity** (Gran, Gutowski, & Papadopoulos, and T.O.): only  $pp$ -waves ( $e^a{}_\mu$ ), strings ( $B_A$ ) and domain walls ( $C, C'$ ).

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$pp$ -waves ( $e^a{}_{\mu}$ ), strings ( $B_A$ ) and domain walls ( $C, C'$ ).

We are going to focus on the domain walls associated to the 3-form  $C'$  since we need to know the associated deformation parameter in order to couple  $C$  to **supergravity** . We consider the **ungauged** theory with only chiral **supermultiplets** and **superpotential**

## 12 – Domain-wall solutions of $N = 1$ supergravity

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The metric of a 4-d **domain-wall solution** can always be written in the form

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If the  $Z^i = Z^i(y)$  the **gravitino Killing spinor** equation  $\delta_\epsilon \psi_\mu = 0$  is solved by

$$(e^{-i\alpha/2} \epsilon) \pm i\gamma^{012} (e^{-i\alpha/2} \epsilon)^* = 0, \quad e^{i\alpha} \equiv \mathcal{L}/|\mathcal{L}|.$$

and  $H(y)$  satisfies the “ **$H$  flow equation**”

$$\partial_{\underline{y}} H^{-1/2} = \pm 2|\mathcal{L}|.$$

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These two first-order **flow equations** imply the second-order **supergravity** equations of motion.

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In the static gauge  $\partial X^\mu / \partial \xi^m = \delta^\mu_m$  it can be seen that this action is invariant to lowest order in fermions under the **supersymmetry** transformations of  $g_{\mu\nu}$ ,  $Z^i$ ,  $C'_{\mu\nu\rho}$  if  $\beta = \pm 1/4$  and the **spinors** satisfy the **BPS domain-wall** projection  $(e^{-i\alpha/2}\epsilon) \pm i\gamma^{012}(e^{-i\alpha/2}\epsilon)^* = 0$ .

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Thus, we consider the bulk **supergravity** action,

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and the **brane** source action

$$S_{\text{brane}} = - \int d^4x \mathbf{f}(\mathbf{y}) \left\{ |\mathcal{L}| \sqrt{|g_{(3)}|} \pm \frac{1}{4!} \epsilon^{mnp} C'_{\underline{mnp}} \right\},$$

where  $\mathbf{f}(\mathbf{y})$  is a distribution function of **domain walls** common transverse direction  $x^3 \equiv \mathbf{y}$ :  $\mathbf{f}(\mathbf{y}) = \delta^{(1)}(\mathbf{y} - \mathbf{y}_0)$  for a single domain wall placed at  $\mathbf{y} = \mathbf{y}_0$  etc.

The equations of motion that follow from  $S \equiv S_{\text{bulk}} + S_{\text{brane}}$  are

$$\mathcal{E}_{\mathbf{g}}^{\mu\nu} = -\frac{\kappa^2}{2} \mathbf{f}(\mathbf{y}) |\mathcal{L}| \frac{\sqrt{|g_{(3)}|}}{\sqrt{|g|}} g_{(3)}^{mn} \delta_m^\mu \delta_n^\nu,$$

$$\mathcal{G}^{ij*} \mathcal{E}_{\mathbf{g}i^*} = -\frac{\kappa^2}{8} \mathbf{f}(\mathbf{y}) \frac{\sqrt{|g_{(3)}|}}{\sqrt{|g|}} e^{i\alpha} \mathcal{N}^i,$$

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The fourth equation is that of  $\mathbf{g}(x)$  and  $\mathbf{y}$  states that  $C'$  is the dual of the **scalar** potential.

## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

It can now be checked that the Einstein and scalar equations of motion are identically satisfied if  $H(y)$  and the scalars  $Z^i(y)$  satisfy the *sourceful flow equations*

$$\partial_{\underline{y}} Z^i = \pm \mathbf{g}(y) e^{i\alpha} \mathcal{N}^i H^{1/2},$$

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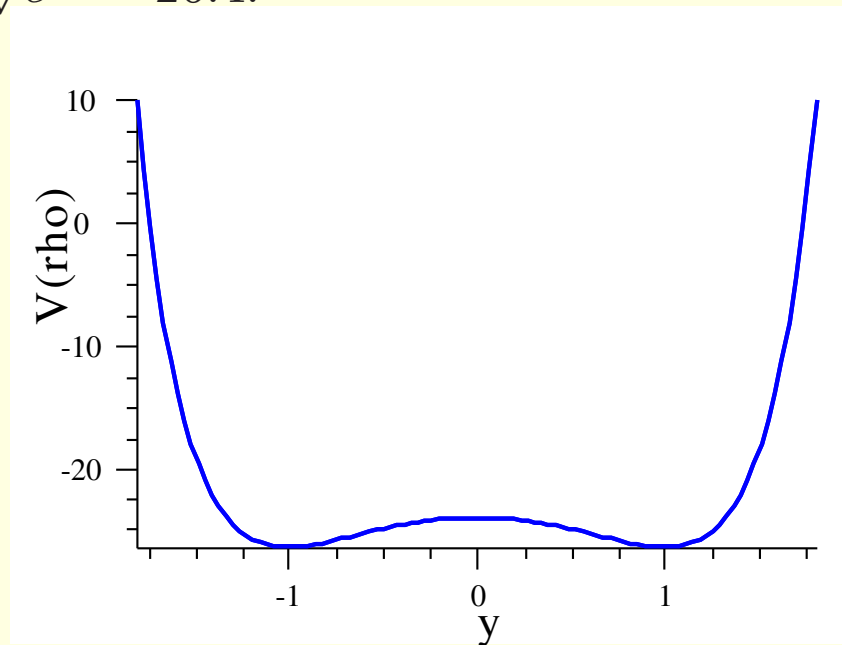
Observe that the space-dependent coupling constant  $\mathbf{g}(x)$ , sourced by **domain wall**, can modify the effective scalar potential **dramatically**.

## 15 – A simple example

Let us consider the model (1 chiral multiplet) defined by

$$\mathcal{K} = |Z|^2, \quad \mathcal{W} = 1, \quad (\mathcal{L} = e^{|Z|^2/2}, \quad \mathcal{N}^Z = 2Z^* e^{|Z|^2/2}).$$

These choices lead to the **Mexican-hat**-type potential  $V = -8(3 - \rho^2)e^{\rho^2/2}$  ( $Z \equiv \rho e^{i\beta}$ ) with a maximum and degenerate minimum at  $\rho = 0$  and  $\rho = +1$  resp. with  $V(0) = -24$ ,  $V(1) = -16\sqrt{e} \sim -26.4$ .



These numbers are irrelevant, as  $V$  will be multiplied by  $\mathbf{g}^2(\mathbf{y})$ , determined by the sources .

The *sourceful flow equations* take the form ( $\beta = \text{Arg } Z = \text{const}$ )

$$\begin{aligned}\partial_{\underline{y}}\rho &= \pm 2\mathbf{g}(\underline{y})\rho e^{\rho^2/2}H^{1/2}, \\ \partial_{\underline{y}}H^{-1/2} &= \pm 2\mathbf{g}(\underline{y})e^{\rho^2/2}.\end{aligned}$$

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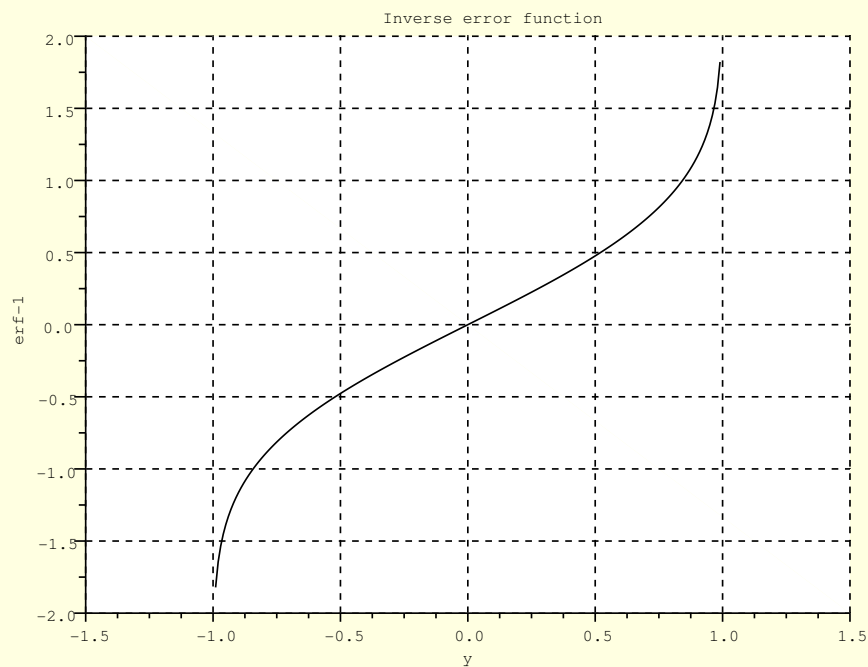
II-b Solutions with  $\mathbf{g} \neq 0$  and  $\partial_{\underline{y}}Z \neq 0$ :

$$H = c/\rho^2,$$

$$\rho = \sqrt{2} \text{erf}^{-1} [\mathbf{G}(\underline{y})], \quad \mathbf{G}(\underline{y}) \equiv \pm \sqrt{\frac{8c}{\pi}} \int \mathbf{g}(\underline{y}) d\underline{y} + d.$$

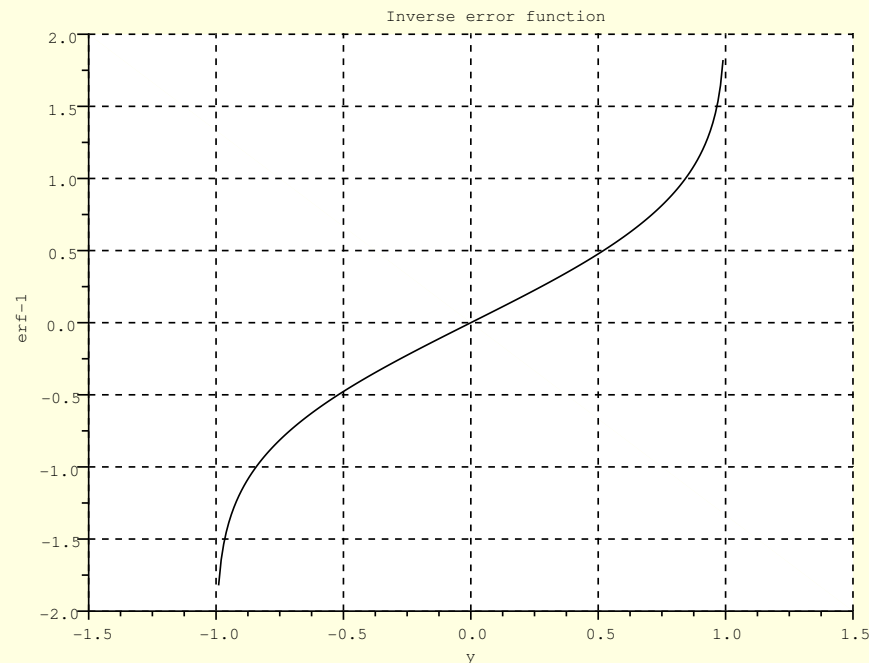
$\operatorname{erf}^{-1}$  is the inverse of the *normalized error function*

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Then, for constant  $\mathbf{g}$ , the solution is restricted to  $y \in (-1, +1)$  and it is locally  $AdS_4$ , but cannot be interpreted as an interpolation between two **vacua unless we cut the spacetime at finite values of  $y$** . To have more general  $\mathbf{g}(y)$  we have to add sources.

## The Tensor Hierarchy of Gauged $N=1, d=4$ Supergravity

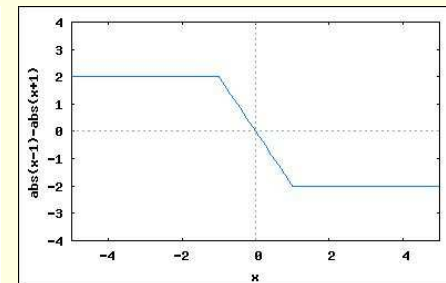
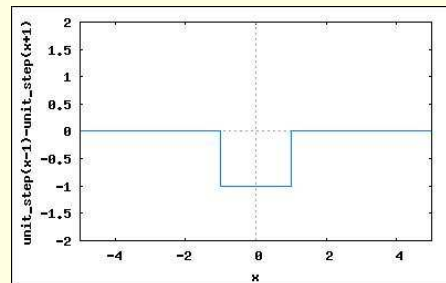
Let us consider, first, a single, infinitely thin **domain-wall** source of tension  $q > 0$  placed at  $y = y_0$ :

$$\mathbf{f}(y) = q\delta(y - y_0), \quad \mathbf{g}(y) = \pm \frac{\kappa^2 q}{16} [\theta(y - y_0) - \theta(y_0 - y)], \quad \mathbf{G}(y) = \frac{\sqrt{c}\kappa^2 q}{\sqrt{32\pi}} |y - y_0| + d.$$

$\mathbf{G}(y)$  is always unbounded and the solution is not well defined.

A possible solution: we introduce two parallel **domain wall** with opposite tension and charge at a different point ( $y = -y_0$  with  $y_0 > 0$  for simplicity) so

$$\begin{aligned} \mathbf{f}(y) &= q\delta(y - y_0) - q\delta(y + y_0), \\ \mathbf{g}(y) &= \pm \frac{\kappa^2 q}{16} [\theta(y - y_0) - \theta(y_0 - y) - \theta(y + y_0) + \theta(-y_0 - y)], \\ \mathbf{G}(y) &= \sqrt{\frac{c}{32\pi}} \kappa^2 q (|y - y_0| - |y + y_0|) + d. \end{aligned}$$



Choosing  $d = \sqrt{\frac{c}{8\pi}} \kappa^2 q y_0$  we can set  $\mathbf{G}(+\infty) = \mathbf{G}(+y_0) = 0$  and  $\rho(y) = \rho(+y_0) = 0 \quad y > y_0$ .

In the interior of the  $\mathbf{g}(y) \neq 0$  region  $\rho$  approaches zero as  $\rho \sim \frac{1}{4} \sqrt{c} \kappa^2 q (y_0 - y)$  so the metric approaches  $AdS_4$

$$H \sim \frac{R^2}{(y_0 - y)^2}, \quad R = \frac{4}{\kappa^2 q}.$$

The value  $\mathbf{G}(-y_0) = \sqrt{\frac{c}{2\pi}} \kappa^2 q y_0 = \mathbf{G}(-\infty)$ , can be tuned by moving the **domain-wall** sources ( $y_0$ ). It has to be smaller or equal than 1.

If  $\mathbf{G}(-y_0) < 1$  then  $\rho(-y_0)$  is finite and  $\rho$  approaches  $y = -y_0$  from the interior of the  $\mathbf{g}(y) \neq 0$  region as

$$\rho \sim -\sqrt{\frac{c}{2\pi}} \frac{\kappa^2 q}{\text{erf}'[\rho(-\infty)/\sqrt{2}]} (y + y_0),$$

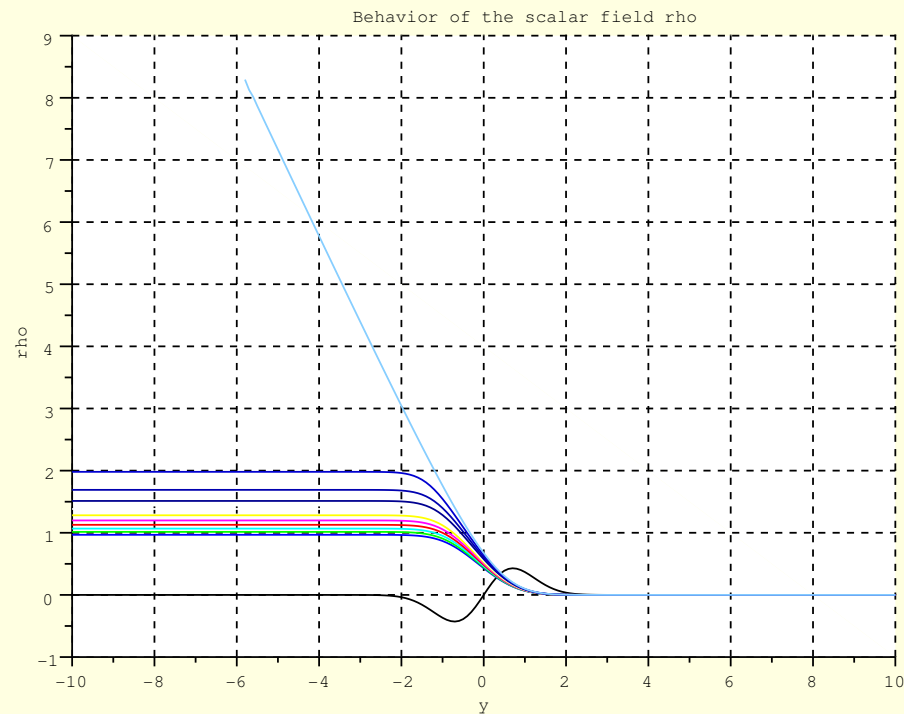
so the metric approaches another  $AdS_4$  region.

This solution we have obtained smoothly interpolates between two  $AdS_4$  regions one of which (the  $\rho = 0$  one) corresponds to a **supersymmetric vacuum** of the theory.

The two **infinitely-thin domain-wall** sources setup can be understood as a crude approximation to the following configuration with **domain-wall** sources of **finite thickness**

$$\mathbf{f}(y) = qye^{-y^2}, \quad \mathbf{g}(y) = \mp \frac{\kappa^2 q}{16} e^{-y^2}, \quad \mathbf{G}(y) = -\frac{\kappa^2 q \sqrt{c}}{8} \text{erf}(y) + d.$$

in which  $\mathbf{g}(y)$  only vanishes asymptotically.



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- ★ We have seen that in some cases **domain-wall** sources have to be introduced to construct sensible solutions. These sources introduce a spacetime-dependent coupling constant  $\mathbf{g}(x)$  that can have dramatic effects.